# The Precession of Mercury's Perihelion 

Where Newton failed and Einstein succeeded.

Dan Wysocki<br>SUNY Oswego<br>QUEST 2014

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- many models were accurate during the lifespan of the people who created them, but over long periods of time began to lose accuracy
- eventually a new model comes along which takes its place, until yet another model replaces that one


## Geocentric Model of the Universe



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- all stars and planets move about the Earth


## Celestial Sphere and Planets



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- the obvious exceptions to this rule are the 7 classical "planets", or "wanderers"
- the Sun, the Moon, Mercury, Venus, Mars, Jupiter, and Saturn
- these planets were thought to circle the Earth in their own paths, presumably closer to Earth than the sphere


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- slowly they would go out of sync with observations


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- some of the brighter stars had drifted by almost an entire degree


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- not a perfect model, but was able to simplify the motions of the planets


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- both cars are moving forward, but the slower car appears to be moving backwards


## Kepler's Laws



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- corrects Copernicus' false assumption that orbits are circular, when they are in fact elliptical


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\mathrm{r}=\mathrm{r}_{0} \frac{1+\varepsilon}{1+\varepsilon \cos \phi} \tag{1}
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- Newton's law of gravity predicts this, while Kepler's laws do not


## Newtonian Precession Predictions

| Planet | $(\delta \dot{\psi})_{\text {obs }}$ | $(\delta \dot{\psi})_{\text {th }}$ |
| :--- | :--- | :--- |
| Mercury | $575.19^{\prime \prime}$ | $532.08^{\prime \prime}$ |
| Venus | $21.6^{\prime \prime}$ | $13.2^{\prime \prime}$ |
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- mystery puzzled physicists for many years


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- a new explanation was in order


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- the planets aren't travelling in a curved path around the Sun, the space around the Sun is itself curved


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- this was one of the first things that Einstein calculated to test his theory out


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- attempts to adapt the model to account for its flaws can have some success, but ultimately are a sign that a new model is needed
- sometimes a completely new approach is needed to be successful


## References I

URL: http: //www. spaceanswers.com/wp-content/uploads/2013/01/Celestial-sphere-illus.jpg.
URL: http://www. astronomy.ohio-state.edu/~pogge/Ast162/Movies/uma.gif.
URL: https://lh3.ggpht.com/cgMuSQK8oEg/TqnJBliUhxI/AAAAAAAAANk/yRe1t5WUbzo/s1600/old+geocentric+model.jpg.
URL: http://www.lasalle.edu/~smithsc/Astronomy/images/marsmovie.gif.
URL: http://faculty.fullerton.edu/cmcconnell/Models/ptolemy.gif.
URL: https://upload.wikimedia.org/wikipedia/commons/0/0e/Cassini_apparent.jpg.
URL: https://upload.wikimedia.org/wikipedia/commons/2/28/Copernican_heliocentrism_diagram2.jpg.

URL: http://faculty.fullerton.edu/cmcconnell/Models/Copernicus_Mars.gif.
URL: https://upload.wikimedia.org/wikipedia/commons/f/f7/An_image_describing_the_semi-major_and_semi-minor_axis_of_eclipse.png.
URL: https://upload.wikimedia.org/wikipedia/commons/thumb/1/1a/Kepler-first-law.svg/500px-Kepler-first-law.svg.png.
URL: https://upload.wikimedia.org/wikipedia/commons/6/69/Kepler-second-law.gif.
URL: https://upload.wikimedia.org/wikipedia/commons/8/89/Precessing_Kepler_orbit_280frames_ e0.6_smaller.gif.
URL: http://www.science4all.org/wp-content/uploads/2013/05/Gravity.jpg.
Carroll, Bradley W. and Dale A. Ostlie. An Introduction to Modern Astrophysics. 2nd ed. Pearson Education, Inc., 2007.

## References II

Fitzpatrick, Richard. Classical Mechanics: an introductory course. 2006. URL: http://web.archive. bibalex. org/web/20060523200517/farside.ph.utexas.edu/teaching/301/lectures/node155.html.

- .Newtonian Dynamics. 2011. URL: http://farside.ph.utexas.edu/teaching/336k/lectures/.

Gerber's Gravity. URL: http://www.mathpages.com/home/kmath527/kmath527.htm.
Halley, Edmund. "Considerations on the Change of the Latitudes of Some of the Principal Fixt Stars". In:
Philosophical Transactions of the Royal Society 30 (1717).
Roseveare, N. T. "Mercury's Perihelion from Le Verrier to Einstein". In: The British Journal for the Philosophy of Science 35 (2), pp. 188-191.
The Precession of the Perihelion of Mercury. URL:
http://www.relativity.li/en/epstein2/read/i0_en/i1_en/.

## Questions?

