

Morphology of the Large and Small Magellanic Clouds using Fundamental Mode Cepheids

Rochester Academy of Science 2014

Daniel Wysocki

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2 Variable Stars

3 Galactic Morphology

Magellanic Clouds

What are they?

Morphology
of LMC and
SMC

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Magellanic
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Galactic
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- two dwarf galaxies which orbit the Milky Way

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- two dwarf galaxies which orbit the Milky Way
- irregular galaxies

Large Magellanic Cloud (LMC)

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- 50kpc away

Small Magellanic Cloud (SMC)

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- 60kpc away

Why are they important?

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- gravitationally interacting with our own galaxy

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- nearby galaxies

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- gravitationally interacting with our own galaxy
- nearby galaxies
 - can be observed in detail

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- gravitationally interacting with our own galaxy
- nearby galaxies
 - can be observed in detail
 - can be used as a distance calibrator to more distant galaxies

Variable Stars

What are they?

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- stars whose luminosity changes with time
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- stars whose luminosity changes with time
- many different types
 - Classical Cepheids, Type II Cepheids, RR Lyrae, MIRA Variables, Delta Scutis, and more

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- the variation can be related to physical properties of the star

- function of a star's brightness over time

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- for periodic variables, time can be transformed into phase

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- for periodic variables, time can be transformed into phase
- shape, amplitude, and period of a star's light curve can reveal many things

- periodic variable stars

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 - period ranges from days to months

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- obey a period-luminosity-color relationship

- luminosity depends on surface area and temperature

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$$\overline{M}_\lambda = \alpha_\lambda \log P + \beta_\lambda + \epsilon_\lambda \quad (2)$$

Period-Luminosity-Color Relationship

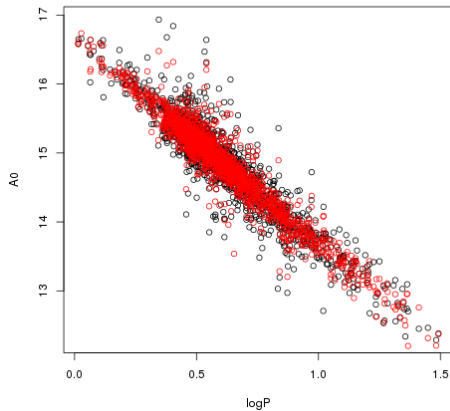
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- imagine a candle, whose luminosity is known

Standard Candles

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- imagine a candle, whose luminosity is known
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- Classical Cepheids can be used in the same way

- difference between apparent and observed magnitudes

$$\mu_i = m_i - M_i \quad (3)$$

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$$\bar{m}_{\lambda,i} = \alpha_{\lambda} \log P_i + \beta_{\lambda} + \mu_i + \epsilon_{\lambda,i} \quad (4)$$

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- for N stars, this results in a system of $2N$ equations with $N + 4$ unknowns

Galactic Morphology

- right ascension (RA or α) is the astronomical equivalent of longitude

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Equatorial Coordinate System

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- introduce distance (D), and the coordinate system now describes three dimensional space

- familiar x, y, z coordinate system

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- can be obtained from equatorial coordinate system through the following transformations

$$x = -D \sin(\alpha - \alpha_0) \cos \delta,$$

$$y = D \sin \delta \cos \delta_0 - D \sin \delta_0 \cos(\alpha - \alpha_0) \cos \delta,$$

$$z = D_0 - D \sin \delta \sin \delta_0 - D \cos \delta_0 \cos \alpha - \alpha_0 \cos \delta$$

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- describe the orientation of a galaxy with respect to Earth

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- can be obtained in different ways

- most common method for obtaining inclination and position angles

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- fits the 2D plane of best fit to the collection of stars

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- fits the 2D plane of best fit to the collection of stars
- basic linear algebra is used to find the 2 angles
- does not do a very good job describing the 3D structure of the galaxy

- less common method for obtaining inclination and position angles

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- fits a 3D ellipsoid to the collection of stars
- principal axis transformation is performed on moment of inertia tensor
- transformation gives eigenvalues and eigenvectors, which are used to describe the size and orientation of the axes of the ellipsoid

- Sukanta Deb, Shashi Kanbur, and H. P. Singh

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- Indo-U.S. Knowledge R&D Joint Networked Center for the Analysis of Variable Star Data

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